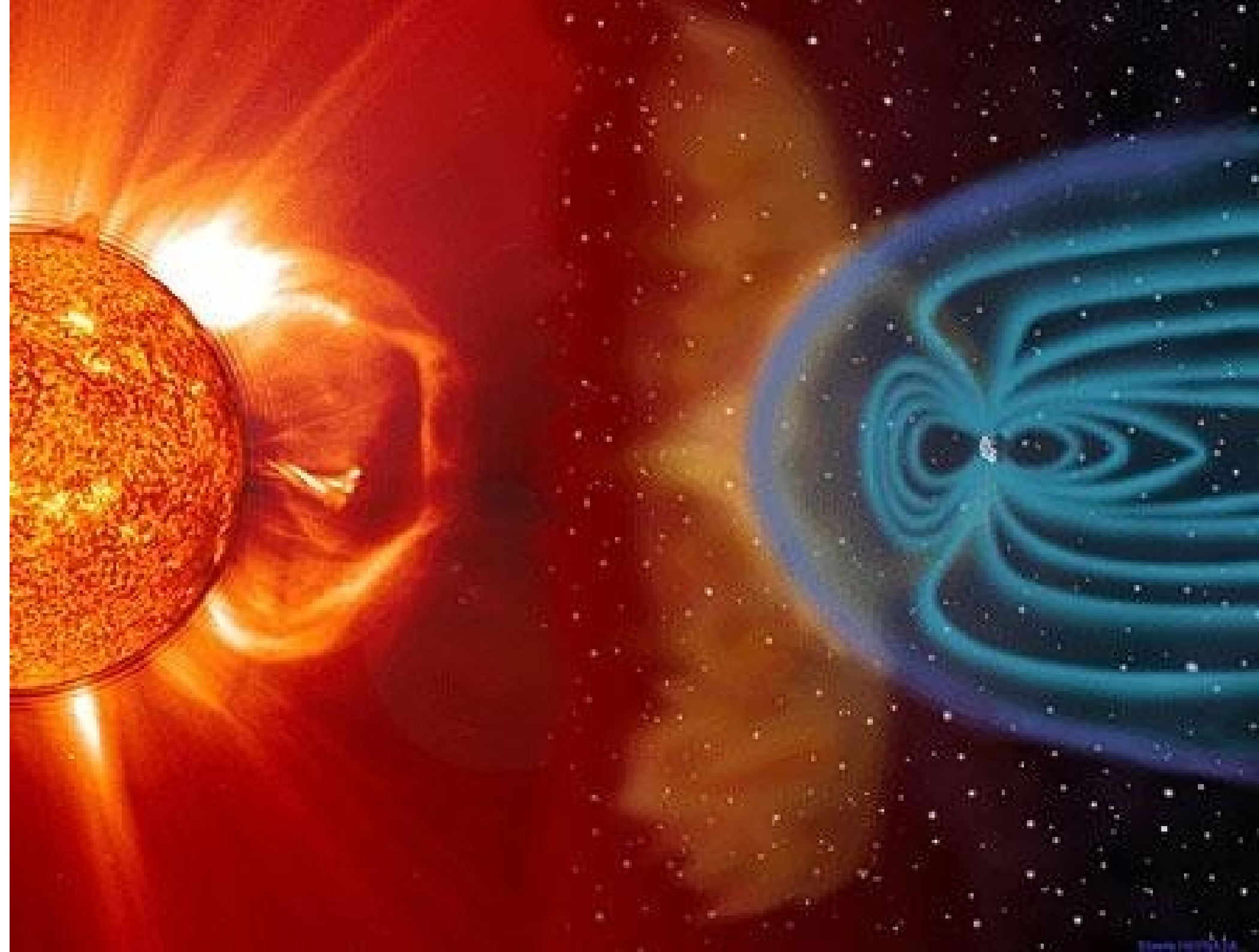
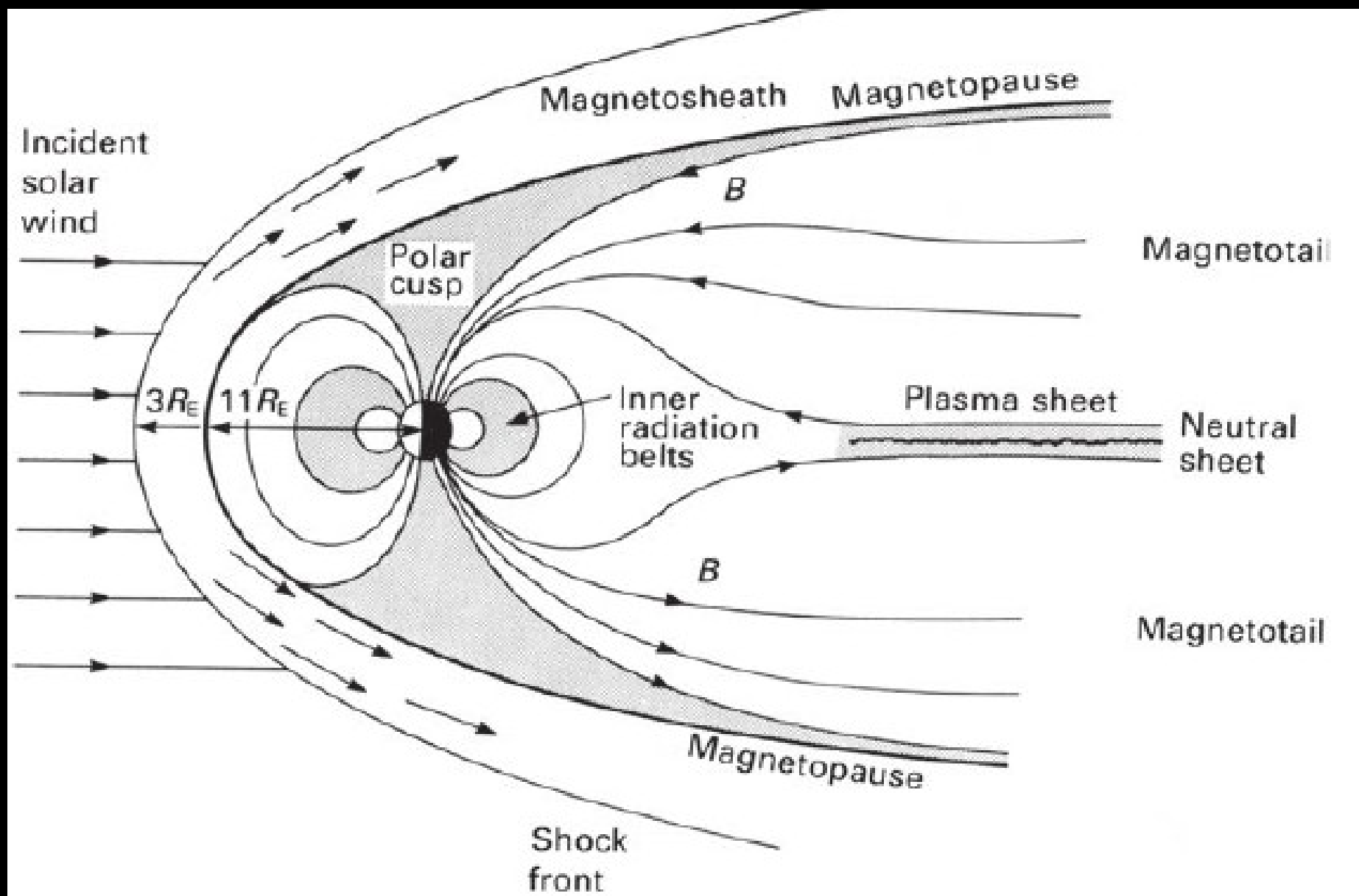
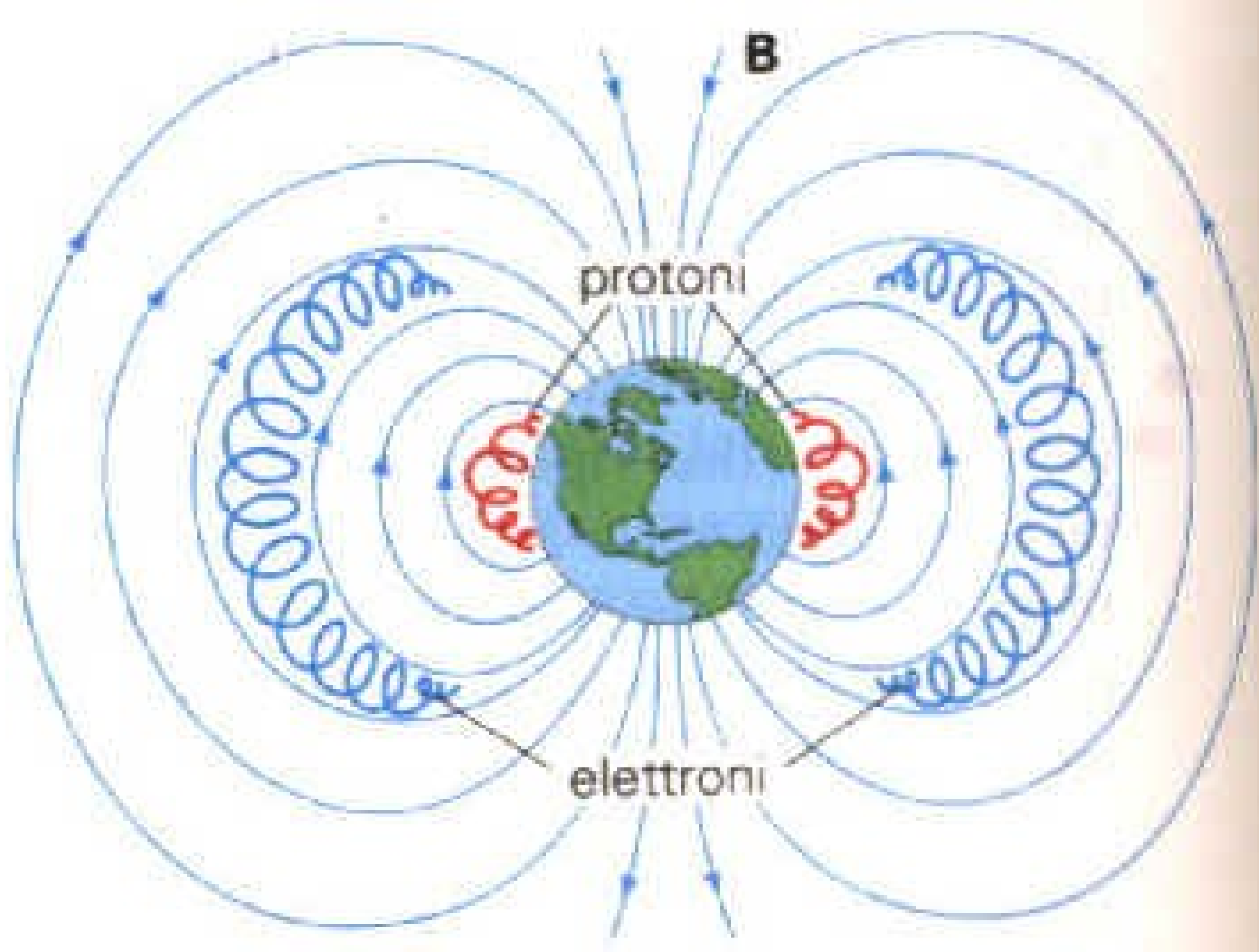


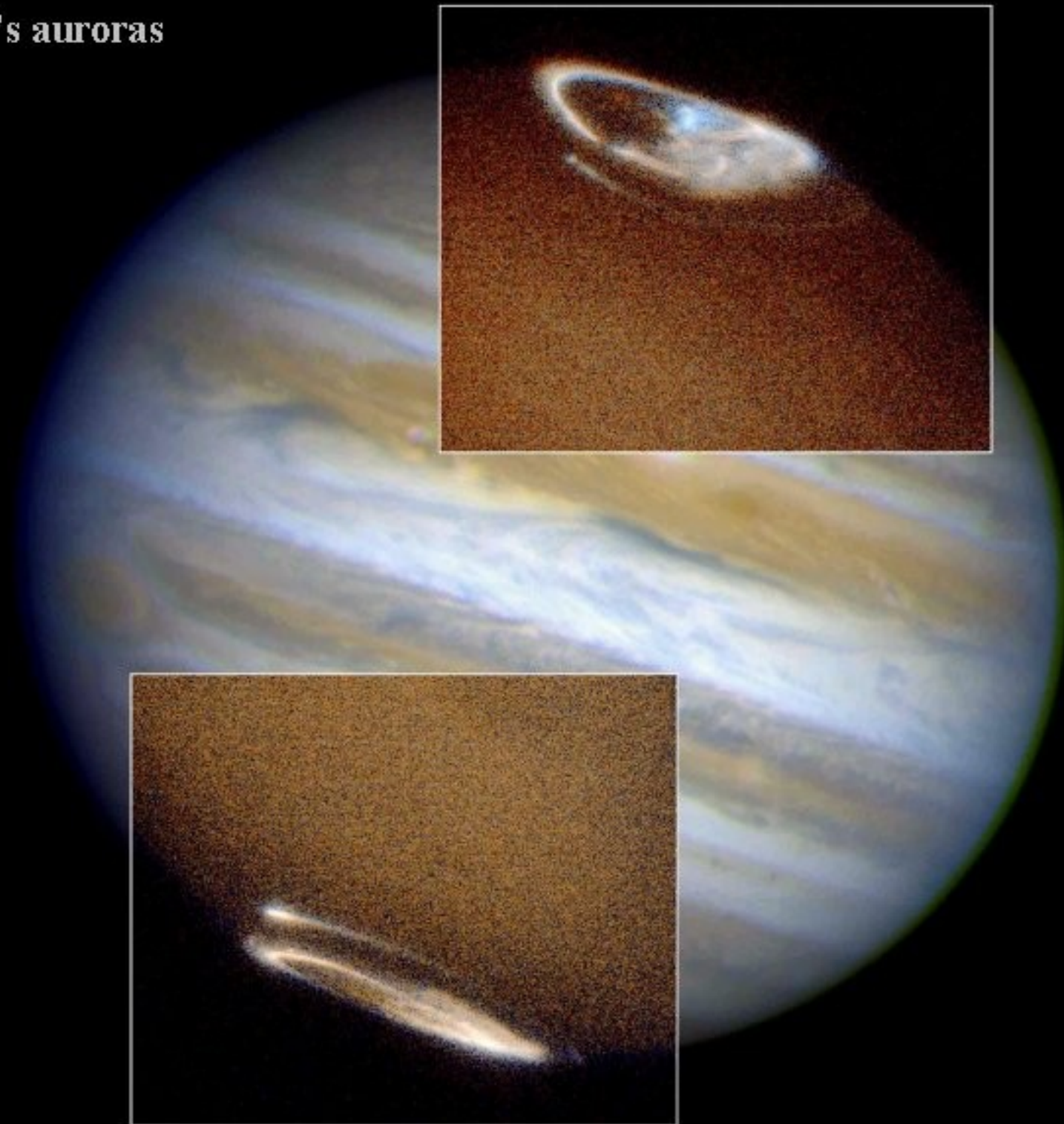
The Solar System

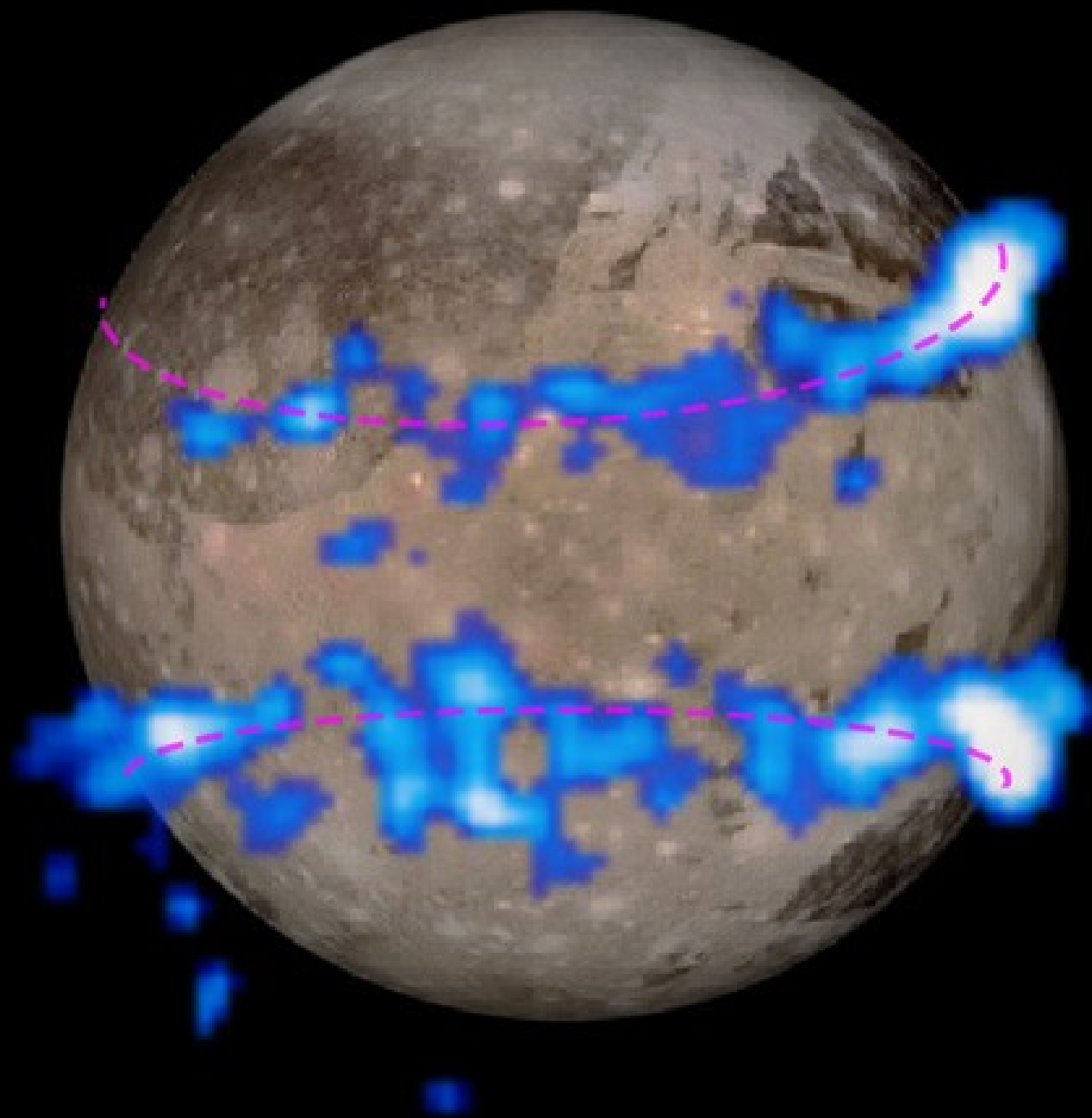






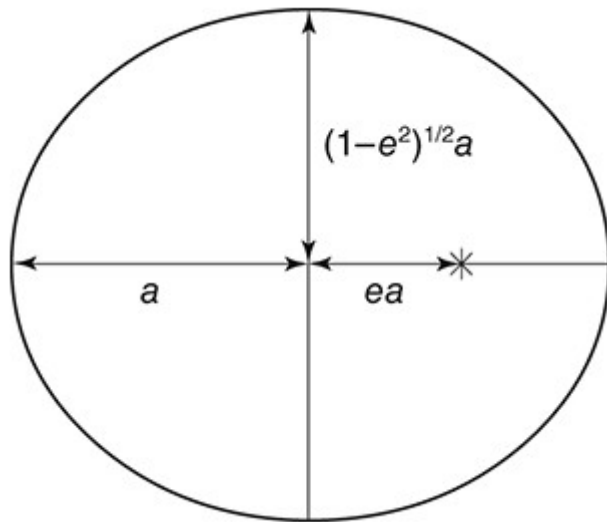
Jupiter's auroras



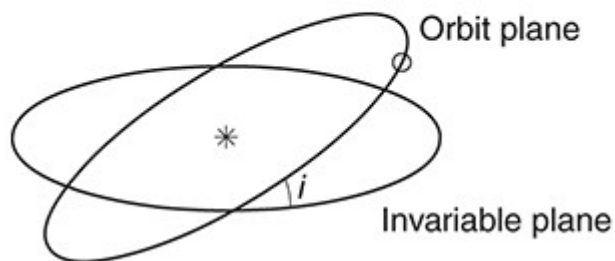


1st Kepler Law

Planetary orbits are elliptical with the sun at a focus.



$$L = \sqrt{1-e^2} \sqrt{GMa} (5.1)$$

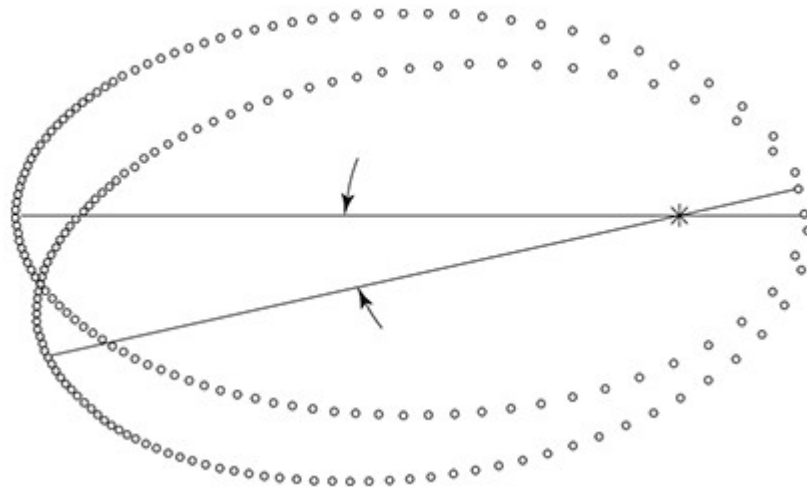


Planet	$M/M_{\oplus}$	a/AU	e	i	Period	$P:P_J$
Mercury	0.055	0.387	0.206	6.34	0.241	1:49.2
Venus	0.815	0.723	0.007	2.19	0.615	1:19.3
Earth	1	1	0.017	1.58	1	1:11.9
Mars	0.108	1.524	0.093	1.67	1.881	1:6.31
Jupiter	317.8	5.203	0.049	0.32	11.86	
Saturn	95.15	9.582	0.056	0.93	29.46	2.48:1
Uranus	14.54	19.19	0.047	1.02	84.02	7.08:1
Neptune	17.15	30.07	0.009	0.72	164.8	13.9:1
Pluto	0.002	39.26	0.245	17.1	247.7	20.9:1

2nd Kepler Law

The radius vector from the sun to a planet sweeps equal areas in equal times.

$$T + U = E_{tot}$$



3rd Kepler Law

The ratio of the square of the period of revolution and the cube of the ellipse semimajor axis is the same for all planets.

$$\text{Grav. acc.} = G M_{\odot} / R^2$$

$$\text{Centrip. acc.} = \Omega^2 R$$

$$\Omega^2 R^3 = G M_{\odot}$$

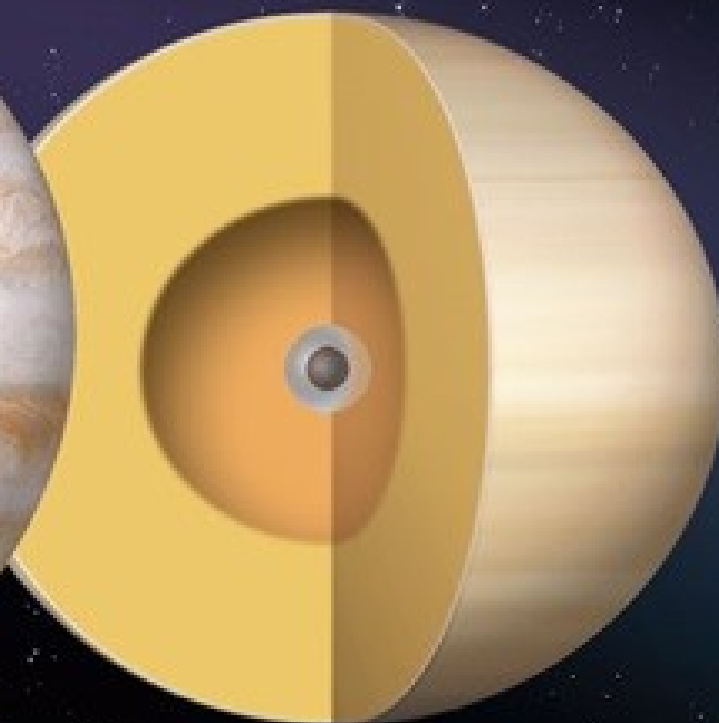


Earth

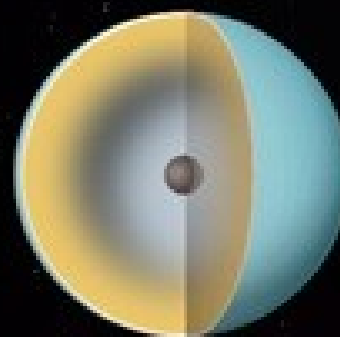
Jupiter



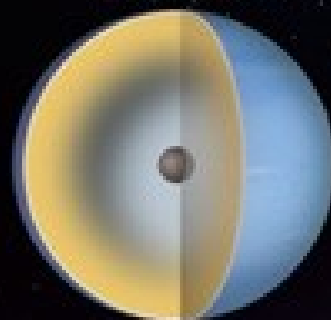
Saturn



Uranus



Neptune

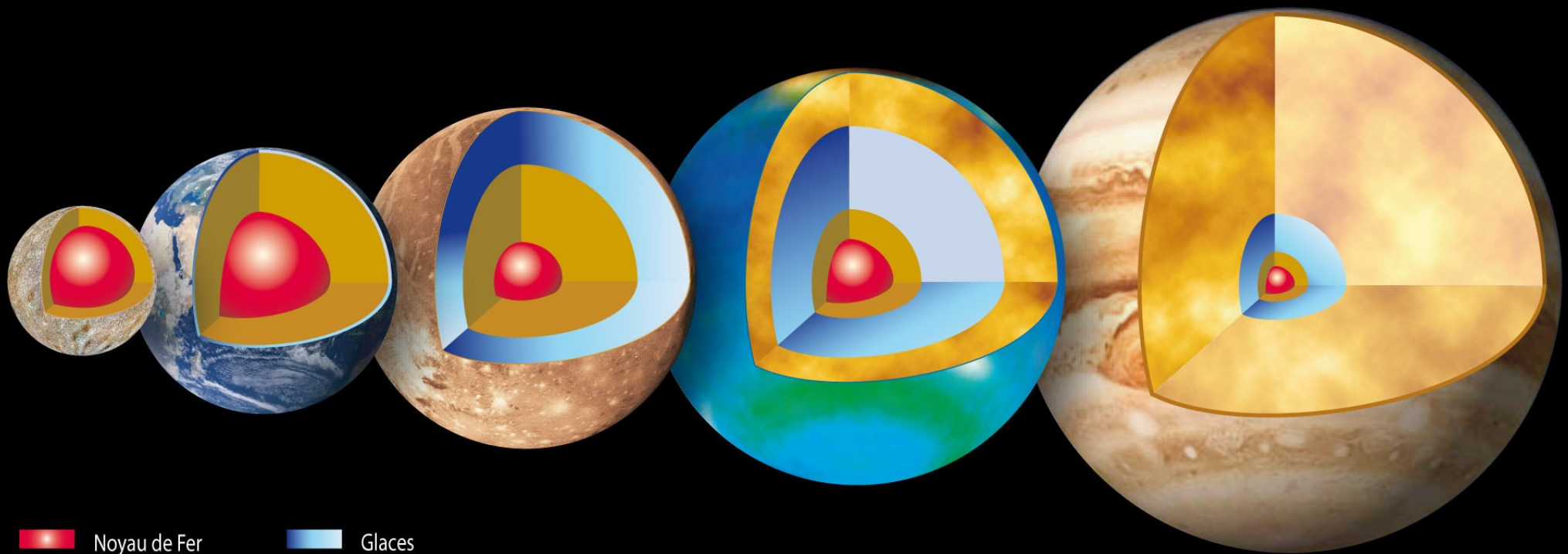


- Rock
- Ionized liquid water, ammonia, and methane
- Liquid metallic hydrogen and atomic helium
- Molecular hydrogen and atomic helium

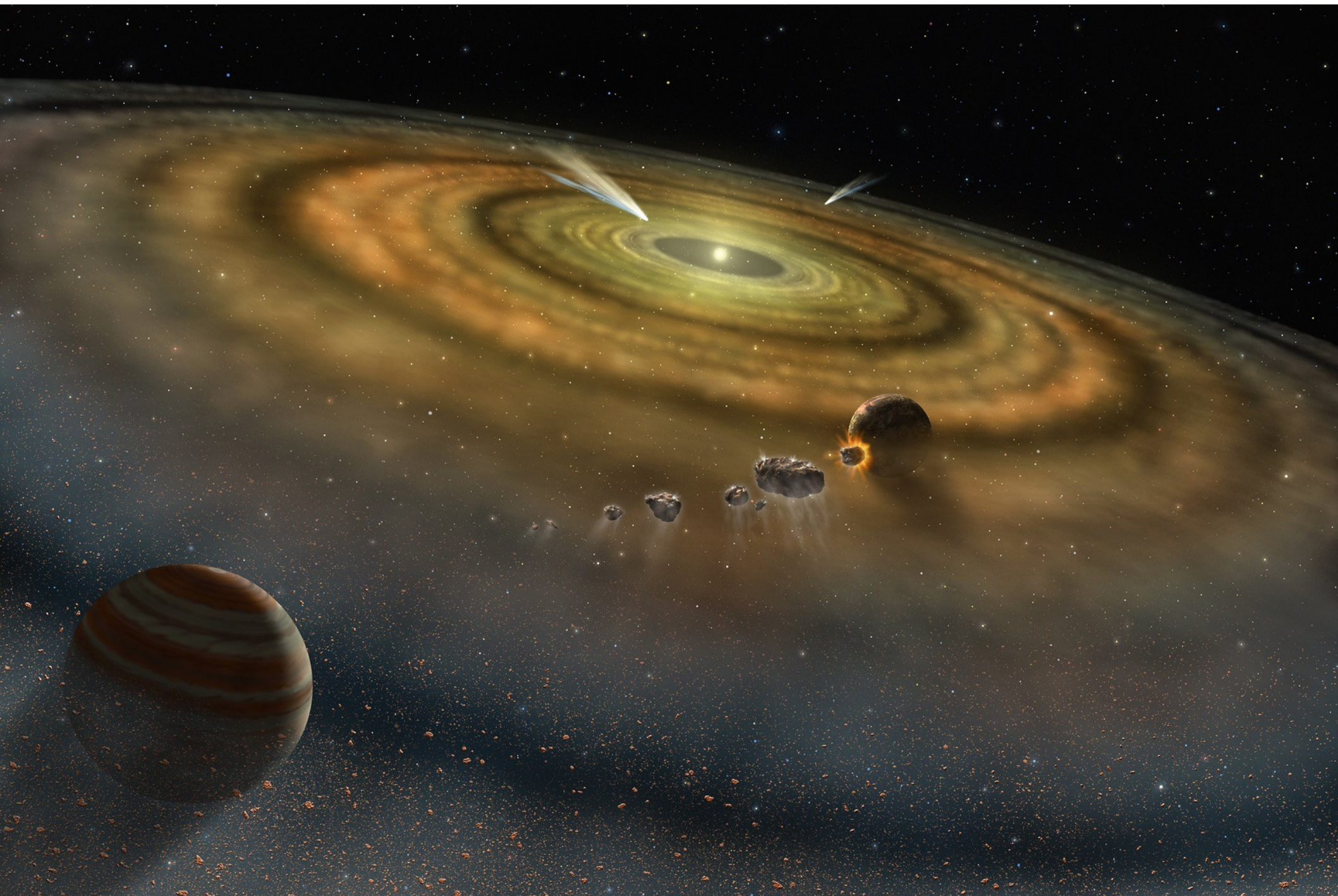
100,000 kilometers

60,000 miles

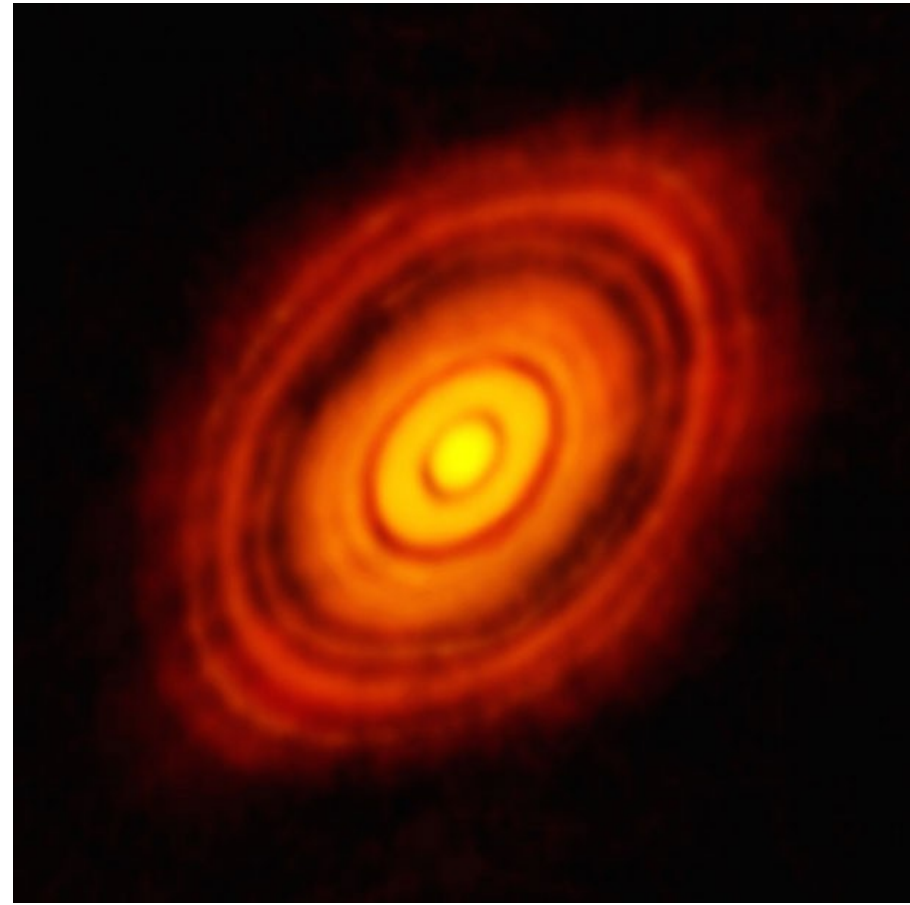
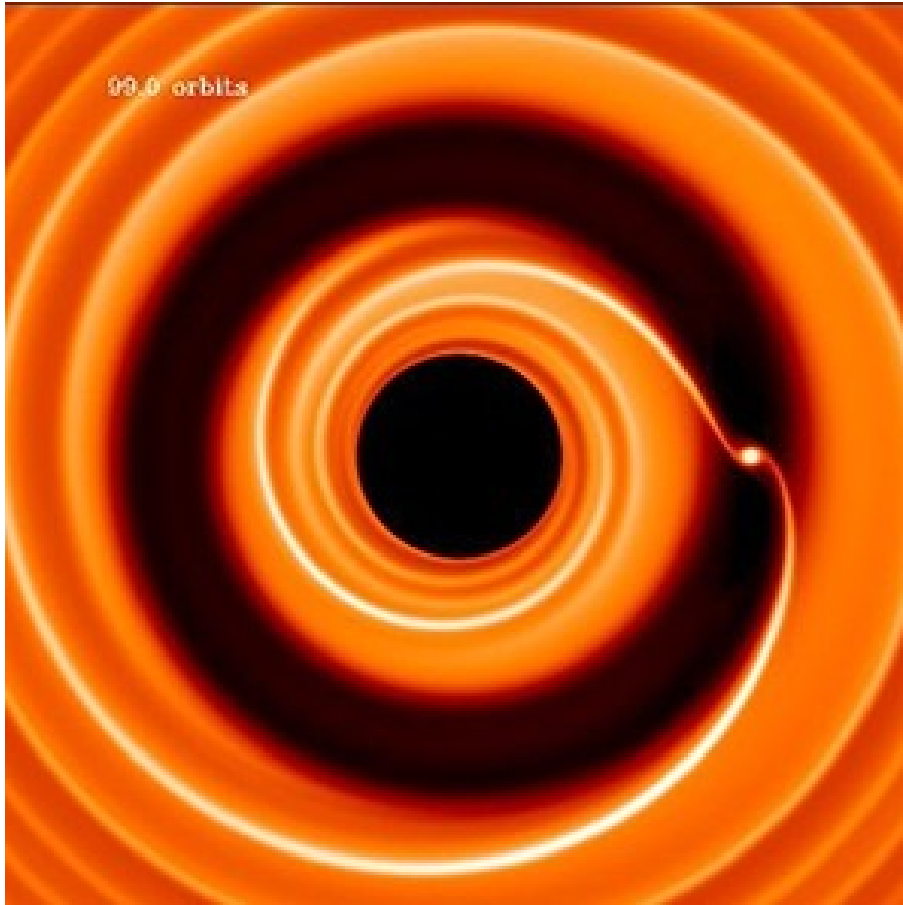




	Noyau de Fer		Glaces
	Manteau silicaté		Hydrogène et hélium



Planetary disk



A planet (right centre) orbiting an (invisible) central star excites a spiral wave in the surrounding gas disc. Through this wave the planet gets angular momentum from the disc inside its orbit, and loses angular momentum to the exterior portion of the disc. These angular- the planet.

Distribution of planets

