

Planck and CMB mapmaking

Andrea Zonca

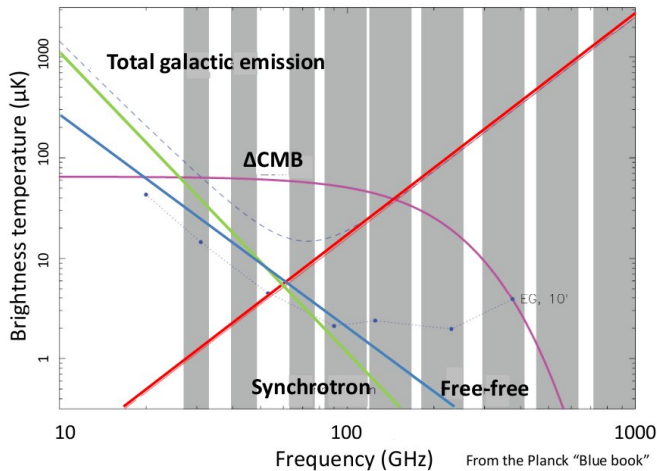
University of California, Santa Barbara

June 10, 2010

- Launch date: 14th May 2009
- Start of first survey: 13th August 2009 (301 days today)
- Start of second survey: 14th February 2009
- 100% sky coverage: 26th April 2010 (LFI) / 28 May 2010 (HFI)
- Mission extended to end of 2011 (30 months)

- Source catalog for followup (ERCSC): December 2010
- End of proprietary period: End of 2012
- First Data release:
 - Calibrated timelines
 - Frequency maps
 - Component maps
 - Source catalogs
 - Cosmological parameters likelihood function
 - Papers

Frequency coverage



Sensitivity

PLANCK	LFI			HFI					
Center freq (GHz)	30	44	70	100	143	217	353	545	857
Angular resolution (FWHM arcmin)	33	24	14	10	7.1	5.0	5.0	5.0	5.0
Sensitivity in I [$\mu\text{K.deg}$] [$\sigma_{\text{pix}} \Omega_{\text{pix}}^{1/2}$]	2.7	2.6	2.6	1.0	0.6	1.0	2.9		
Sensitivity in Q or U [$\mu\text{K.deg}$] [$\sigma_{\text{pix}} \Omega_{\text{pix}}^{1/2}$]	4.5	4.6	4.6	1.8	1.4	2.4	7.3		

From the Planck "Blue book"

WMAP center freq.	23	33	41	61	94
Angular resolution (FWHM arcmin)	49	37	29	20	12.6
Sensitivity in I [$\mu\text{K.deg}$], 1 yr (8 yr)	12.6 (4.5)	12.9 (4.6)	13.3 (4.7)	15.6 (5.5)	15.0 (5.3)

Planck satellite

Planck

Current status

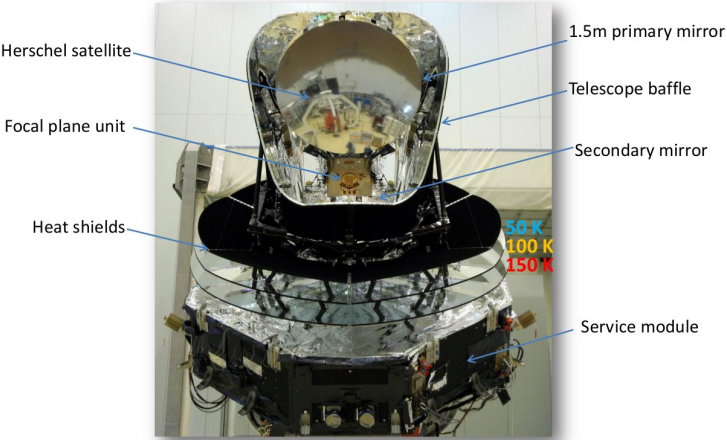
Instruments and
cooling

Published results

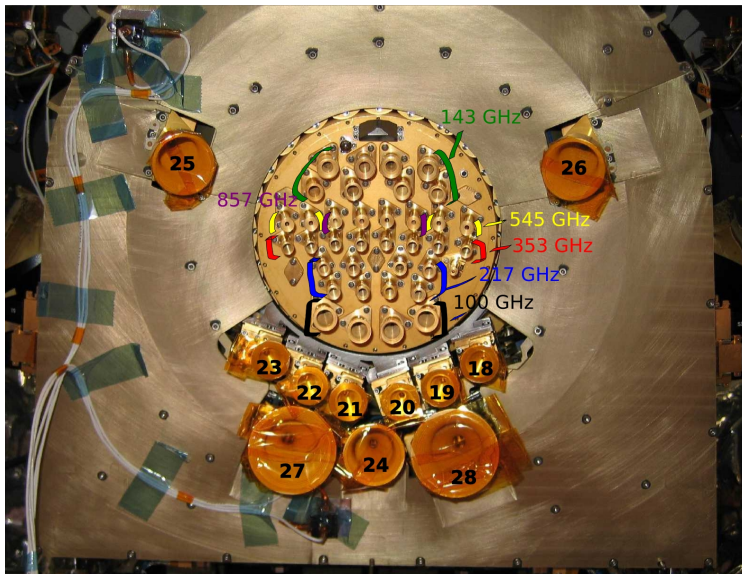
CMB Mapmaking

Scanning

Destriping



Planck focal plane



Planck

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Planck focal plane



Planck

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Low Frequency Instrument radiometer

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Current status

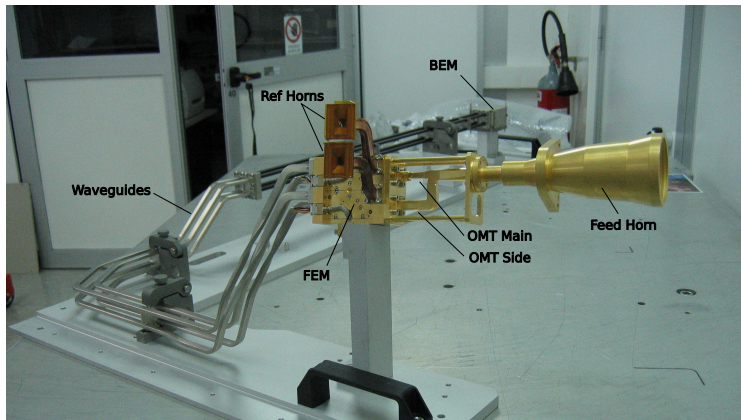
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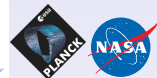
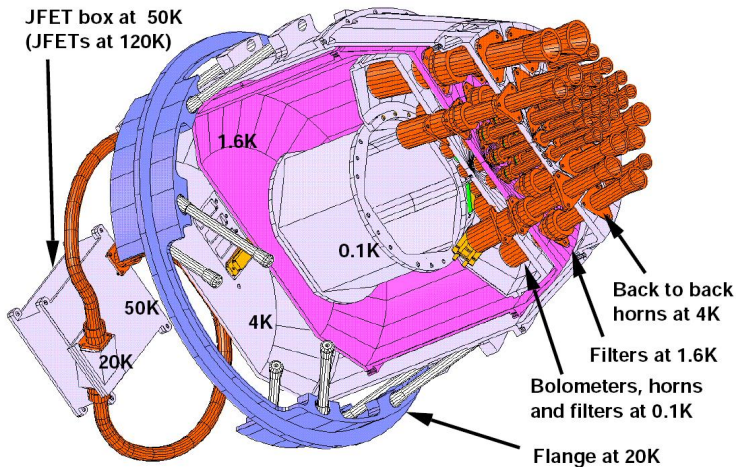
Scanning

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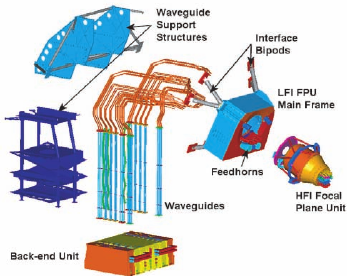
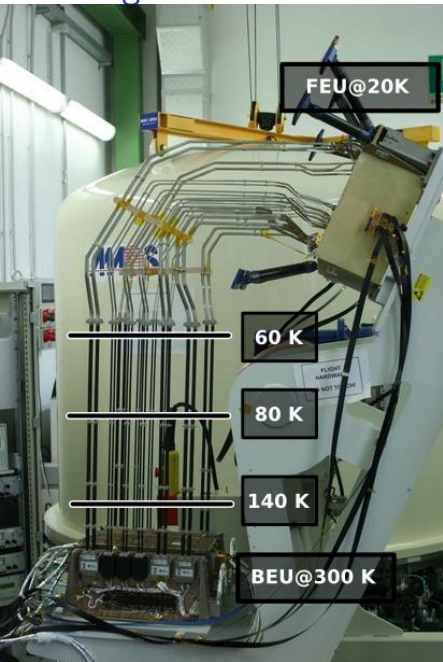


HFI Bolometers array

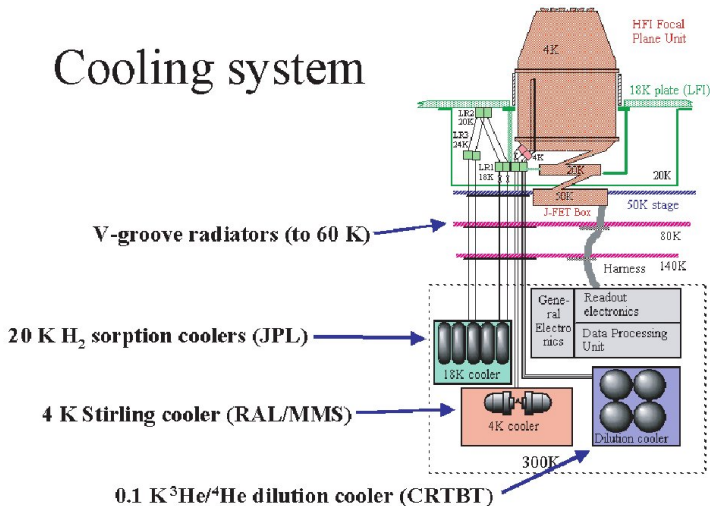
- Planck
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- Published results
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- Scanning
- Destriping



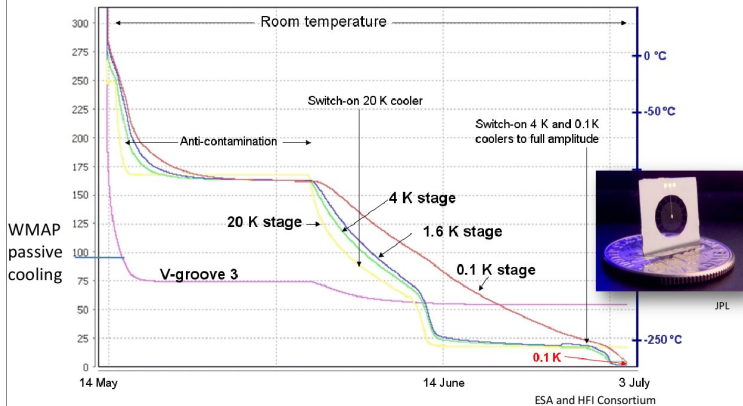
Cooling chain



Cooling system



Inflight cooling profile

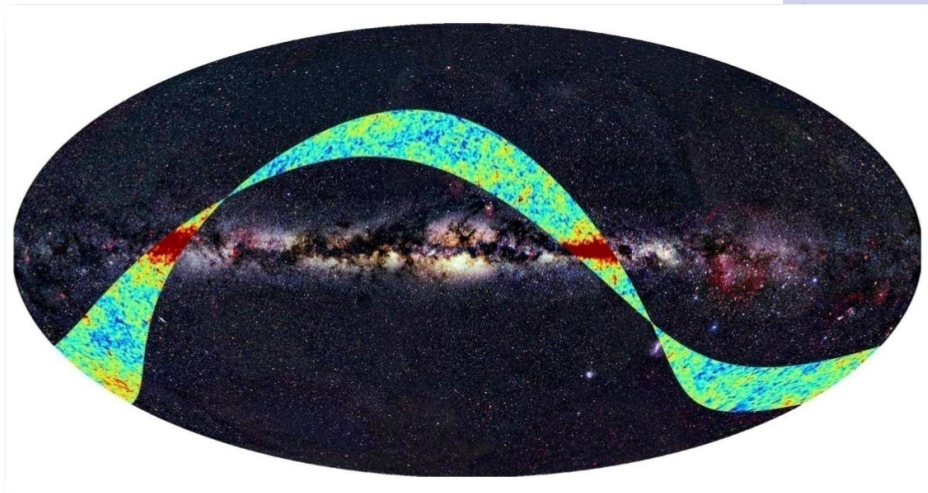


First light survey

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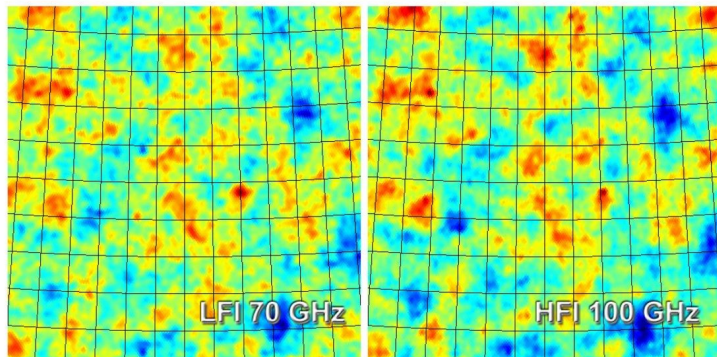
LFI 70 GHz and HFI 100 GHz patch

Planck

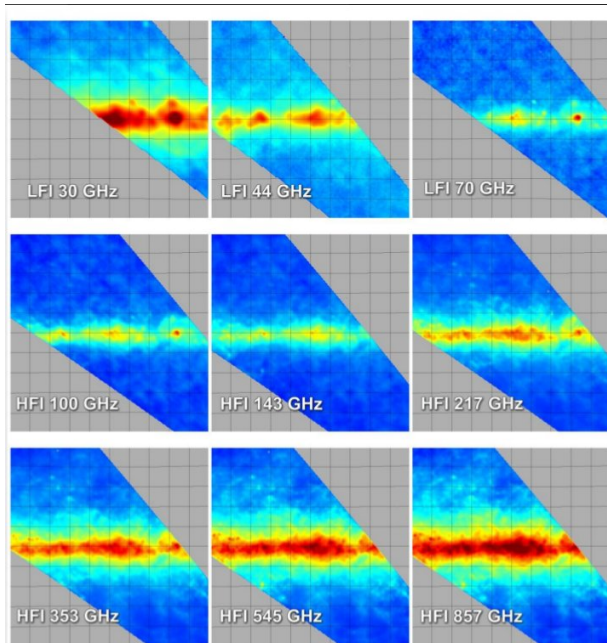
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20°x20° patches around the galaxy



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CMB Mapmaking

Scanning

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Cold dust 12-70K 545/857 GHz

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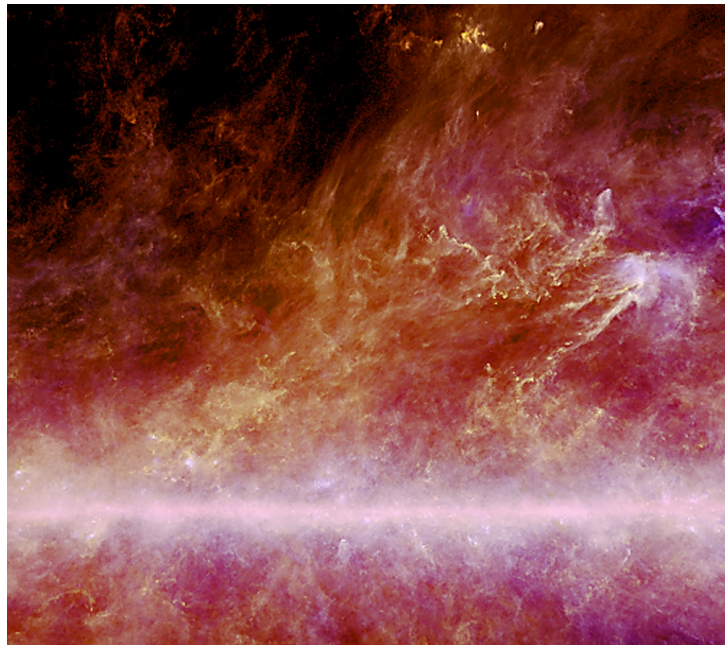
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CMB Mapmaking

Scanning
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Perseus, Planck (30/353/857) vs DSS

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Orbit and scanning

Videos:

PLANCK_LAGRANGE_PC_03.wmv

moll_G_ring.mov

- Optimal maximum likelihood mapmaker
- Destriper

$$y = Pm_{in} + Fa + w \quad (1)$$

y : radiometer output

P : pointing matrix

m_{in} : input map

a : baselines

F : baselines to TOD

w : white noise TOD

Destriping in time domain

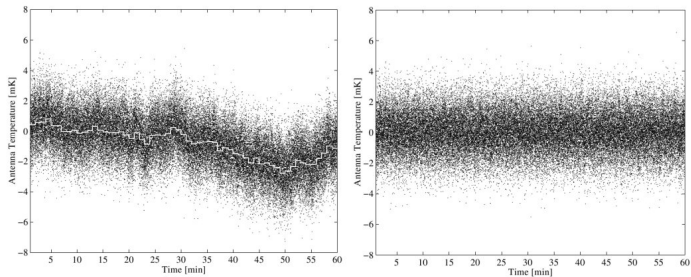


Figure by T. Poutanen

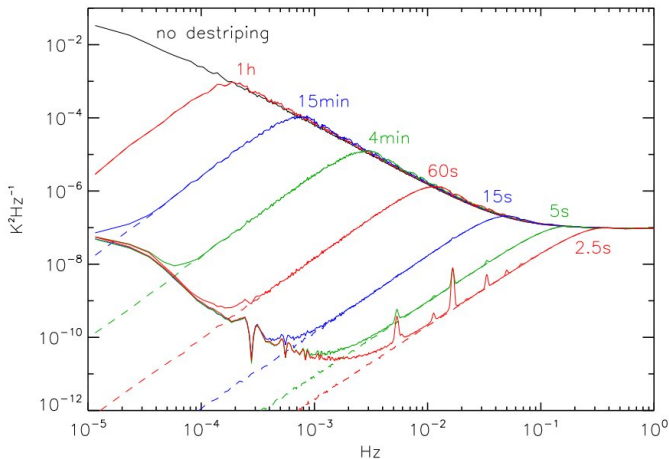
Destriping in frequency domain

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CMB Mapmaking

Scanning
Destriping



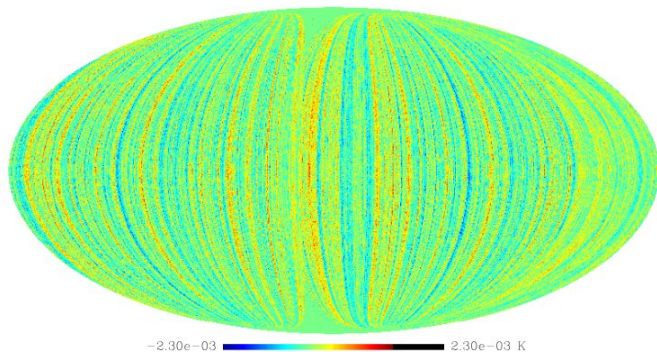
Simulated coadded map

Planck

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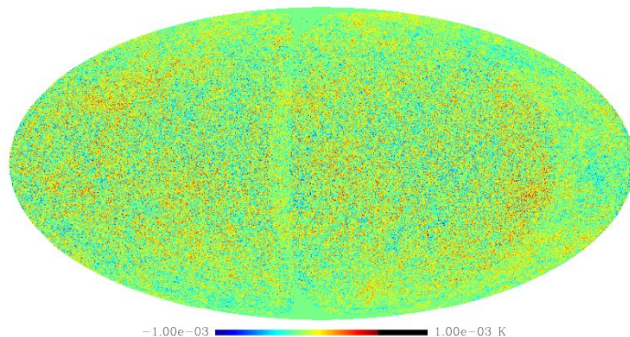
Simulated destriped map 7 degrees

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Simulated destriped map 28 degrees

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