

# The galactic cocktail

a blind approach to the  
component separation

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# the cocktail-party problem

- two people speaking simultaneously:  $s_1, s_2$
- two microphones recording time signals:  $x_1, x_2$

$$\begin{aligned}x_1(t) &= a_{11}s_1 + a_{12}s_2 \\x_2(t) &= a_{21}s_1 + a_{22}s_2\end{aligned} \quad \rightarrow \mathbf{x} = \mathbf{A}\mathbf{s}$$

**A** = mixing matrix

having  $x_1, x_2$ , how to get  $s_1, s_2$  ?

# the problem in astrophysics

- superposition of  $N$  different emissions

$$\tilde{x}(\mathbf{r}, \nu) = \sum_{j=1}^N \tilde{s}_j(\mathbf{r}, \nu)$$

- experiment with  $M$  frequency channels

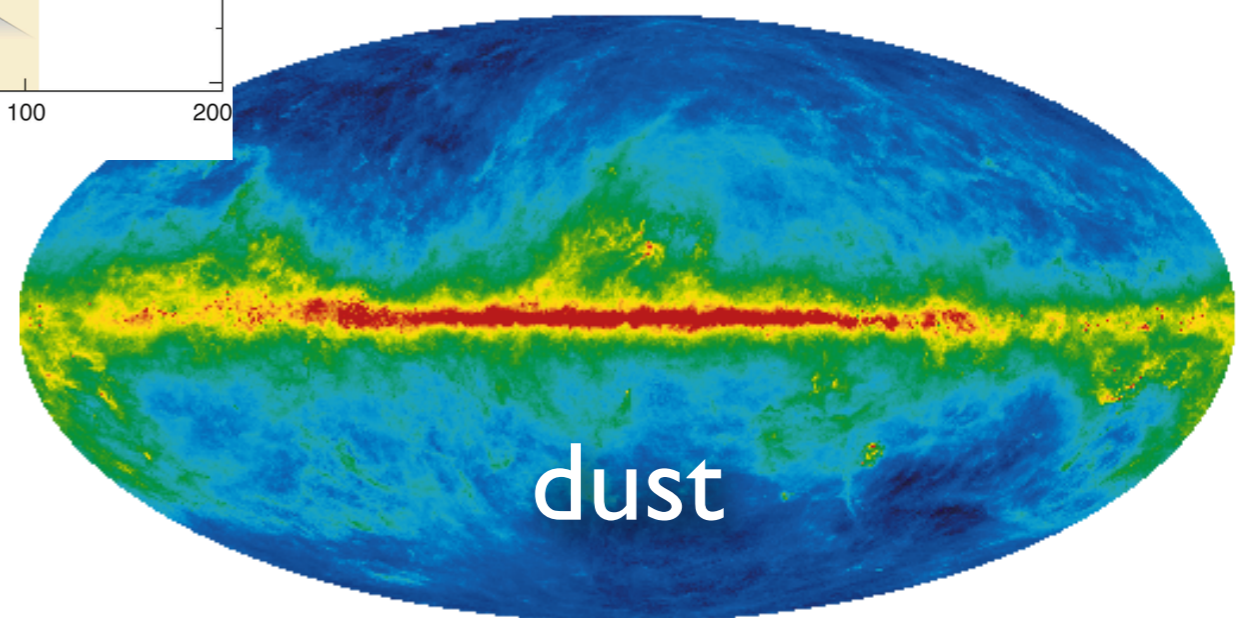
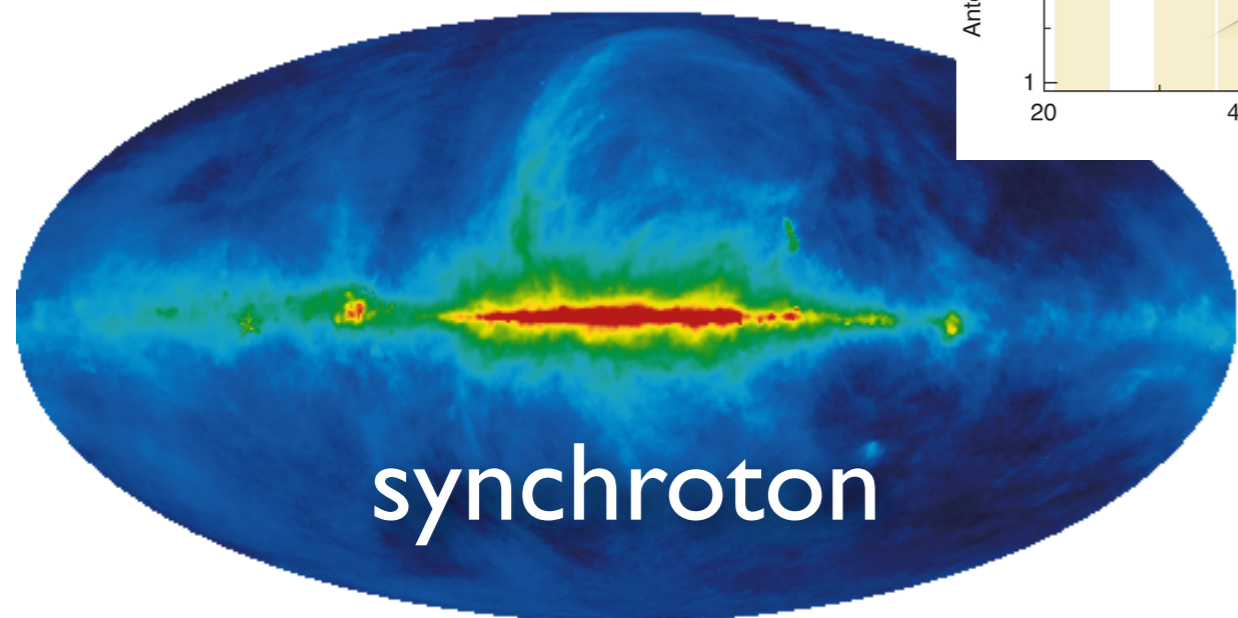
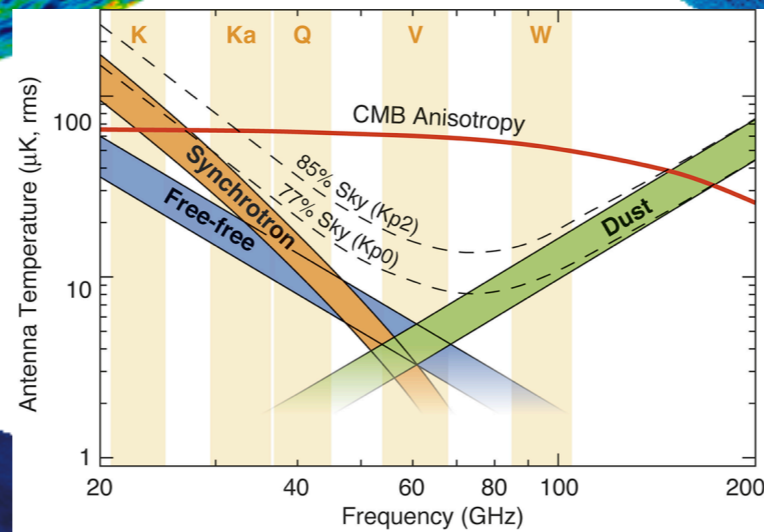
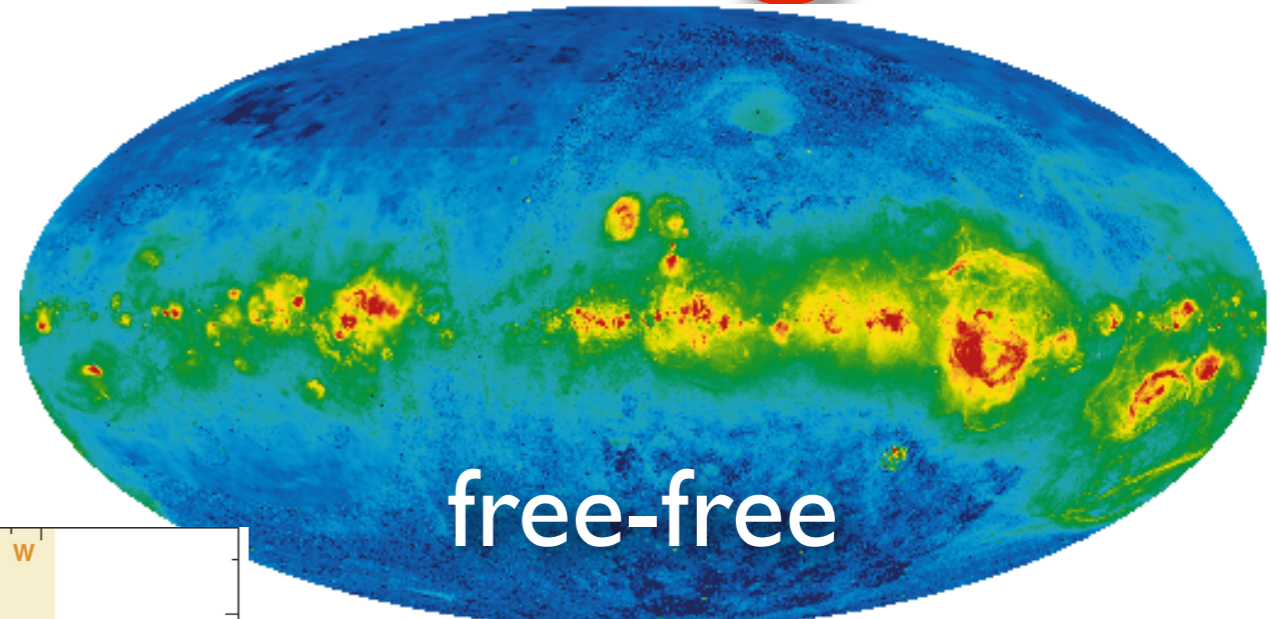
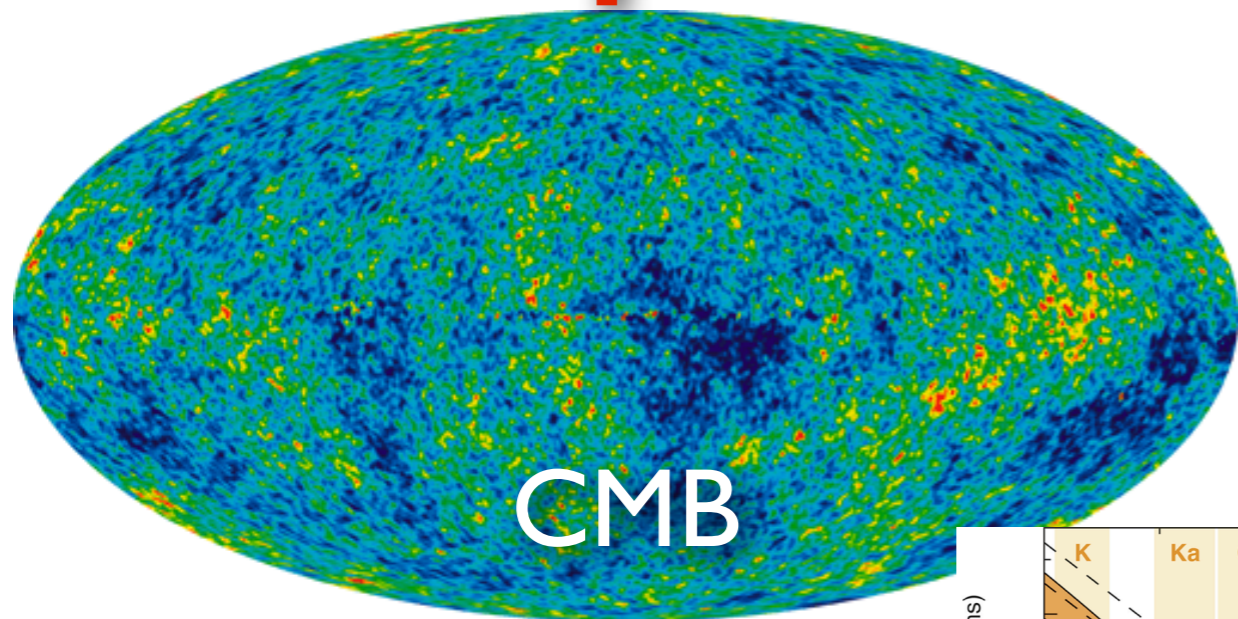
$$\mathbf{x} = \mathbf{A}\mathbf{s} \quad (+ \text{ noise})$$

mixing matrix  $M \times N$  characterized by PSF and frequency response

*how to separate the different emissions?*

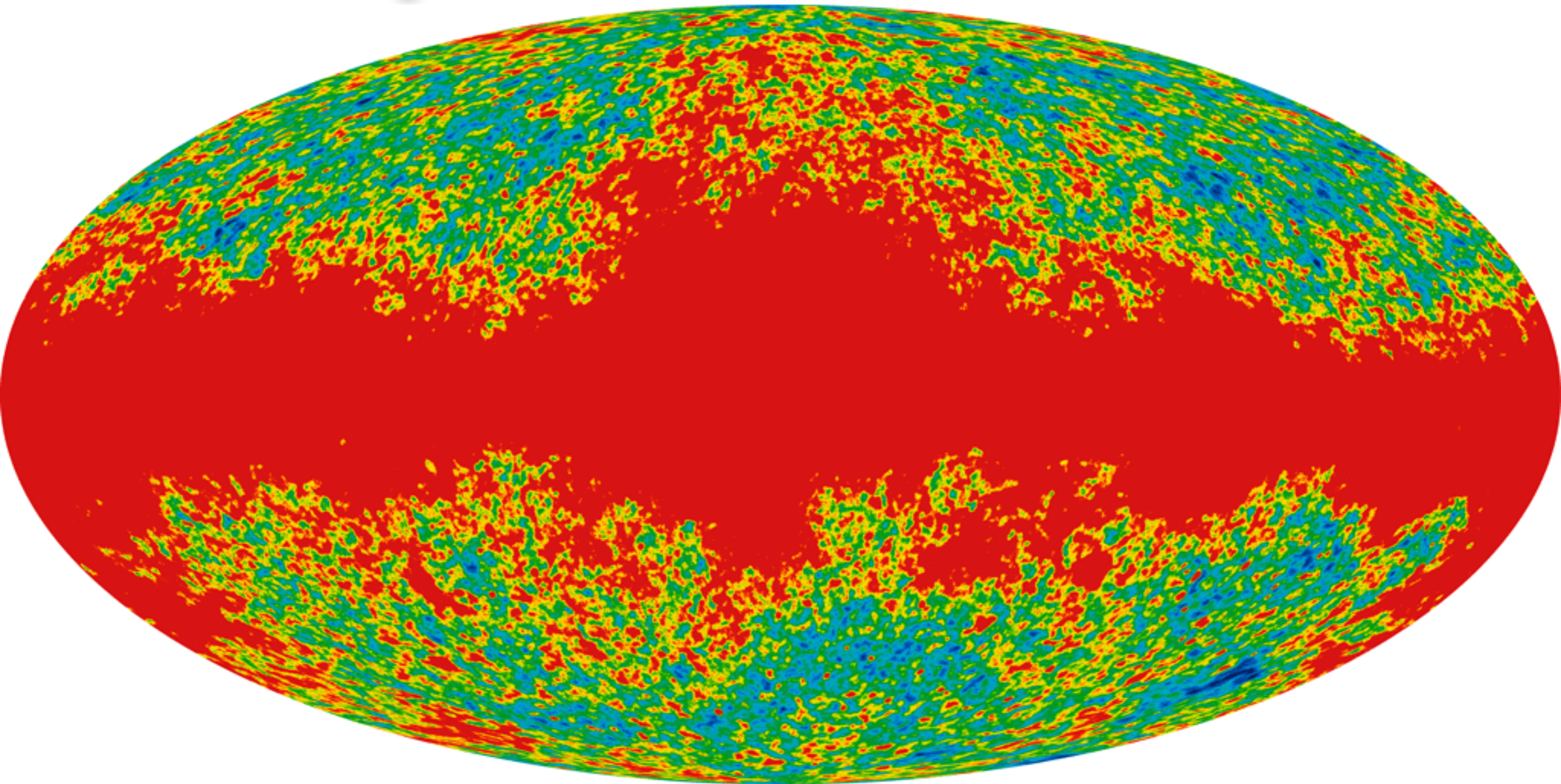


# example: microwaves signals





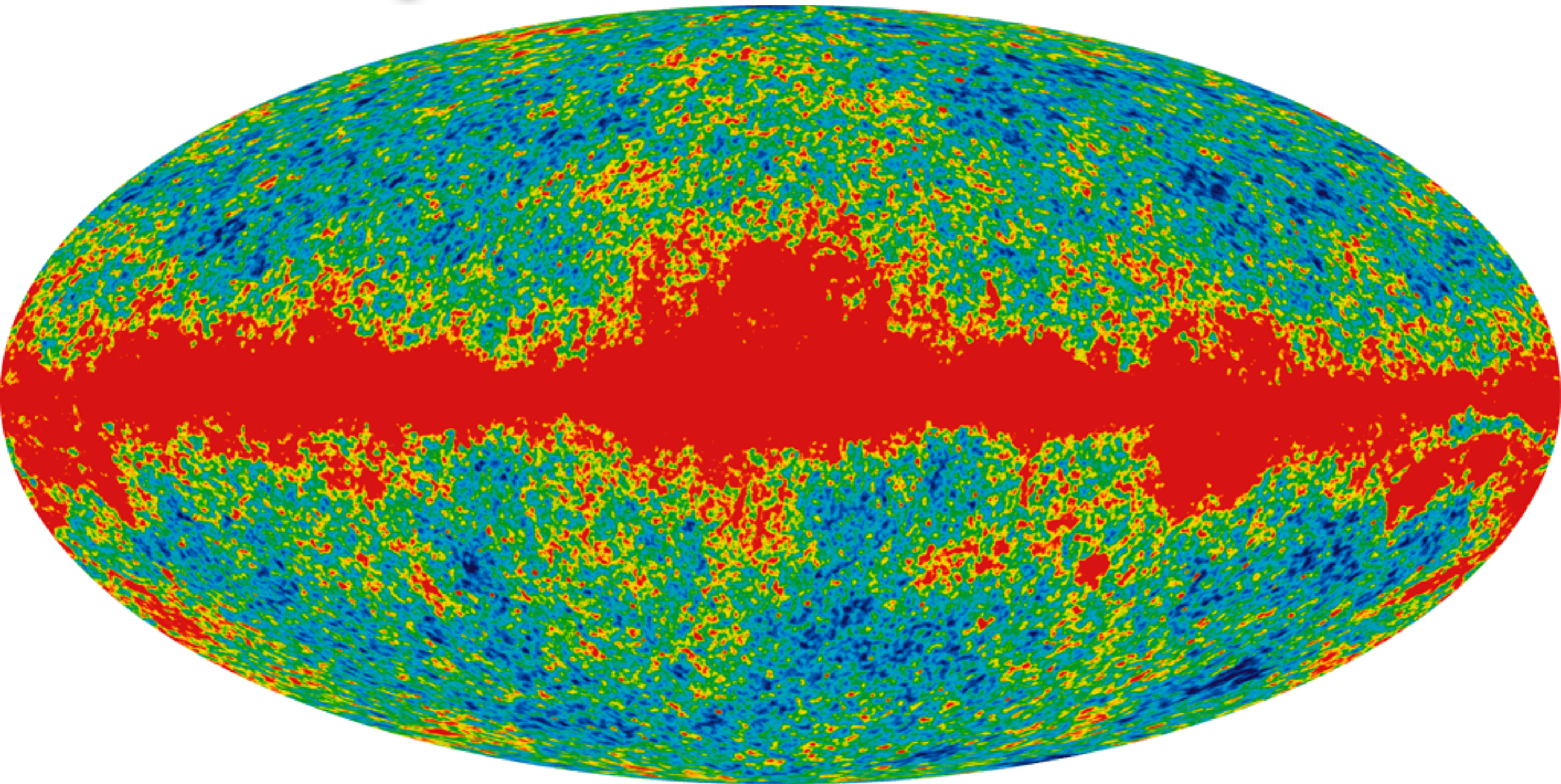
# example: microwaves data



WMAP 7yr - K 23 GHz



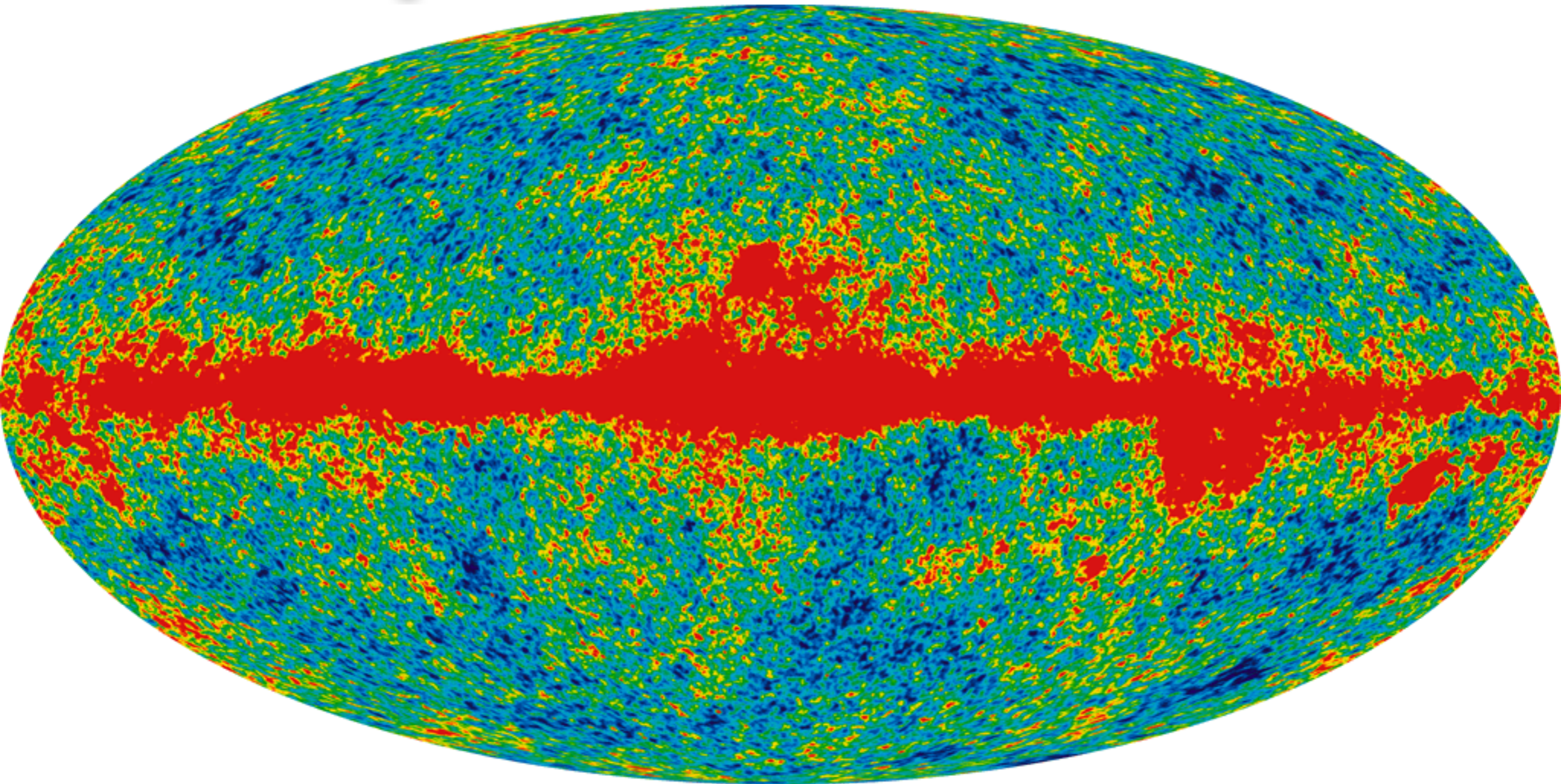
# example: microwaves data



WMAP 7yr - Ka 33 GHz



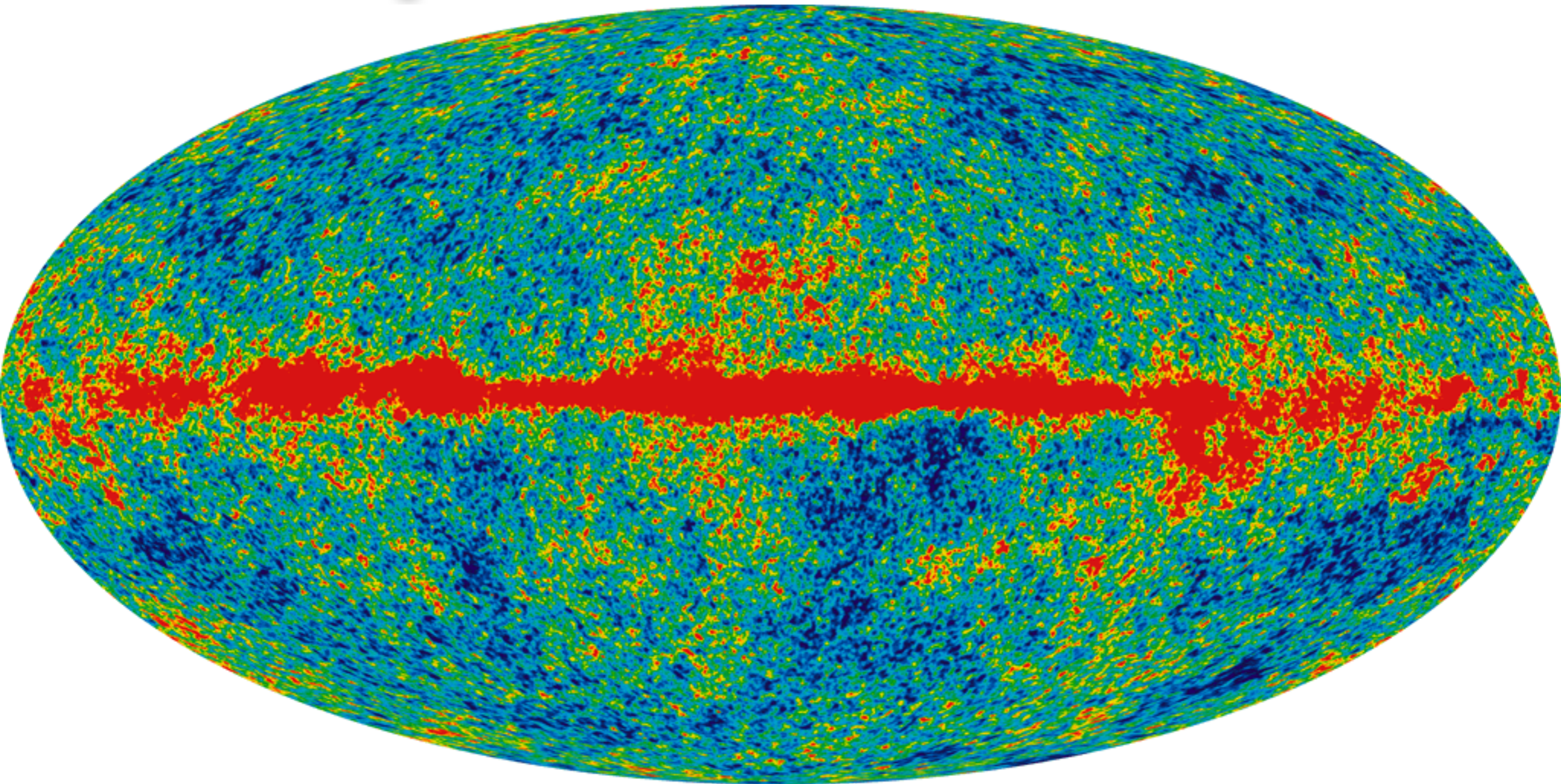
# example: microwaves data



WMAP 7yr - Q 41 GHz



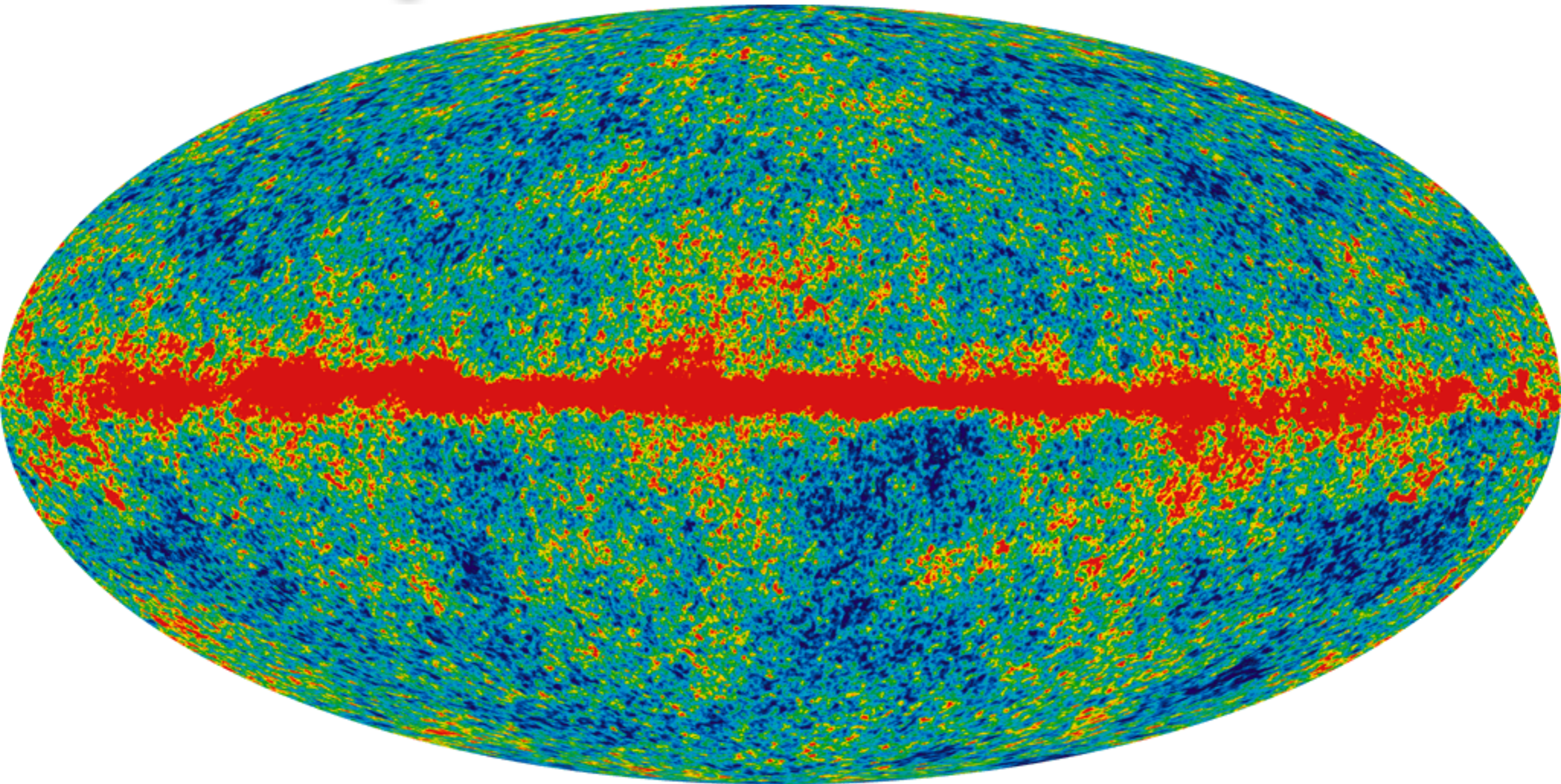
# example: microwaves data



WMAP 7yr - V 61 GHz



# example: microwaves data



WMAP 7yr - W 94 GHz



# ICA: a blind approach

two assumptions:

- the signals  $s$  are statistical independent
- all the signals, but at most one, have a non-Gaussian distribution



the **I**ndependent **C**omponent **A**nalysis can solve the component separation problem without ‘a priori’ knowledge on the signals



# principle of ICA

solution: find a transformation  $\mathbf{W}$  such that the vector  $\mathbf{y}=\mathbf{W}\mathbf{x}$  has independent components

*we need a measure of independence!*

Central Limit Theorem:

the distribution of a sum of random independent variables tends towards a gaussian distribution



# non-Gaussian is independent

$$y = \mathbf{w}^T \mathbf{x} = \mathbf{w}^T \mathbf{A} \mathbf{s} = \mathbf{z}^T \mathbf{s}$$

- $y$  linear combination of  $s_i \Leftrightarrow$  more gaussian
- $y$  becomes least gaussian when  $y = s_i \quad (\mathbf{z} = z_i)$
- $\mathbf{w}^T$  maximizes the non-gaussianity of  $\mathbf{w}^T \mathbf{x}$



# measure of non-gaussianity

*entropy* of a random variable = degree of information needed to describe the variable, i.e. the coding length (more random, i.e. more unpredictable = larger entropy)

differential entropy  $H$  of a random vector  $\mathbf{y}$  with

density  $f(\mathbf{y})$ : 
$$H(\mathbf{y}) = - \int f(\mathbf{y}) \log f(\mathbf{y}) d\mathbf{y}$$

*a gaussian variable has the largest entropy among all random variables of equal variance [information theory]*

negentropy: 
$$J(\mathbf{y}) = H(\mathbf{y}_{gauss}) - H(\mathbf{y})$$



# negentropy

if  $\langle y \rangle = 0$  and  $\langle y^2 \rangle = 1$  then

$$J(y) \propto [E\{G(y)\} - E\{G(\nu)\}]^2$$

where

- $E$  expected value
- $\nu$  Gaussian variable of zero mean and unit variance
- $G$  non quadratic functions



# preprocessing for ICA

- centering:

$$\tilde{\mathbf{x}} = \mathbf{x} - E\{\mathbf{x}\} \quad (\text{zero mean})$$

- (quasi) whitening:

we want  $\hat{\mathbf{x}}$  white (uncorrelated component and unit variance)

$$\hat{\Sigma} = (\mathbf{C} - \Sigma)^{-1/2} \Sigma (\mathbf{C} - \Sigma)^{-1/2}$$

$$\hat{\mathbf{x}} = (\mathbf{C} - \Sigma)^{-1/2} \tilde{\mathbf{x}}$$

where  $\mathbf{C}$  = data covariance,  $\Sigma$  = noise covariance

⇒ mixing matrix  $\mathbf{A}$  is orthogonal:  $M(M-1)/2$  d.o.f.



# the FastICA algorithm

$\mathbf{W}$  estimated *row by row*, i.e. components *one by one*:

1. choose initial vector  $\mathbf{w}$

$$2. \mathbf{w}_{new} = E\{\hat{\mathbf{x}}g(\mathbf{w}^T \hat{\mathbf{x}})\} - (I + \hat{\Sigma})\mathbf{w}E\{g'(\mathbf{w}^T \hat{\mathbf{x}})\}$$

$$3. \mathbf{w}_{new} = \mathbf{w}_{new} / \|\mathbf{w}_{new}\|$$

4. if not converged back to 2., otherwise new row

- $g=G'$  can be  $p(u) = u^3$ ,  $g(u) = u \exp(-u^2)$ ,  $t(u) = \tanh(u)$

- $(k+1)$ th row must be orthogonal to first  $k$  rows:

orthogonalization step between 2. and 3.

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# recovered components

- linear combination

$$y_j = \sum_{i=1}^M w_{ji} x_{\nu_i} \quad j = 1, \dots, M$$

- frequency scaling

$$(W^{-1})_{\nu_j} / (W^{-1})_{\nu'_j}$$

$$\Rightarrow \beta = \log[(W^{-1})_{\nu_j} / (W^{-1})_{\nu'_j}] / \log(\nu / \nu')$$

- ambiguities:
  - normalization factor
  - extraction order



# applications of FastICA

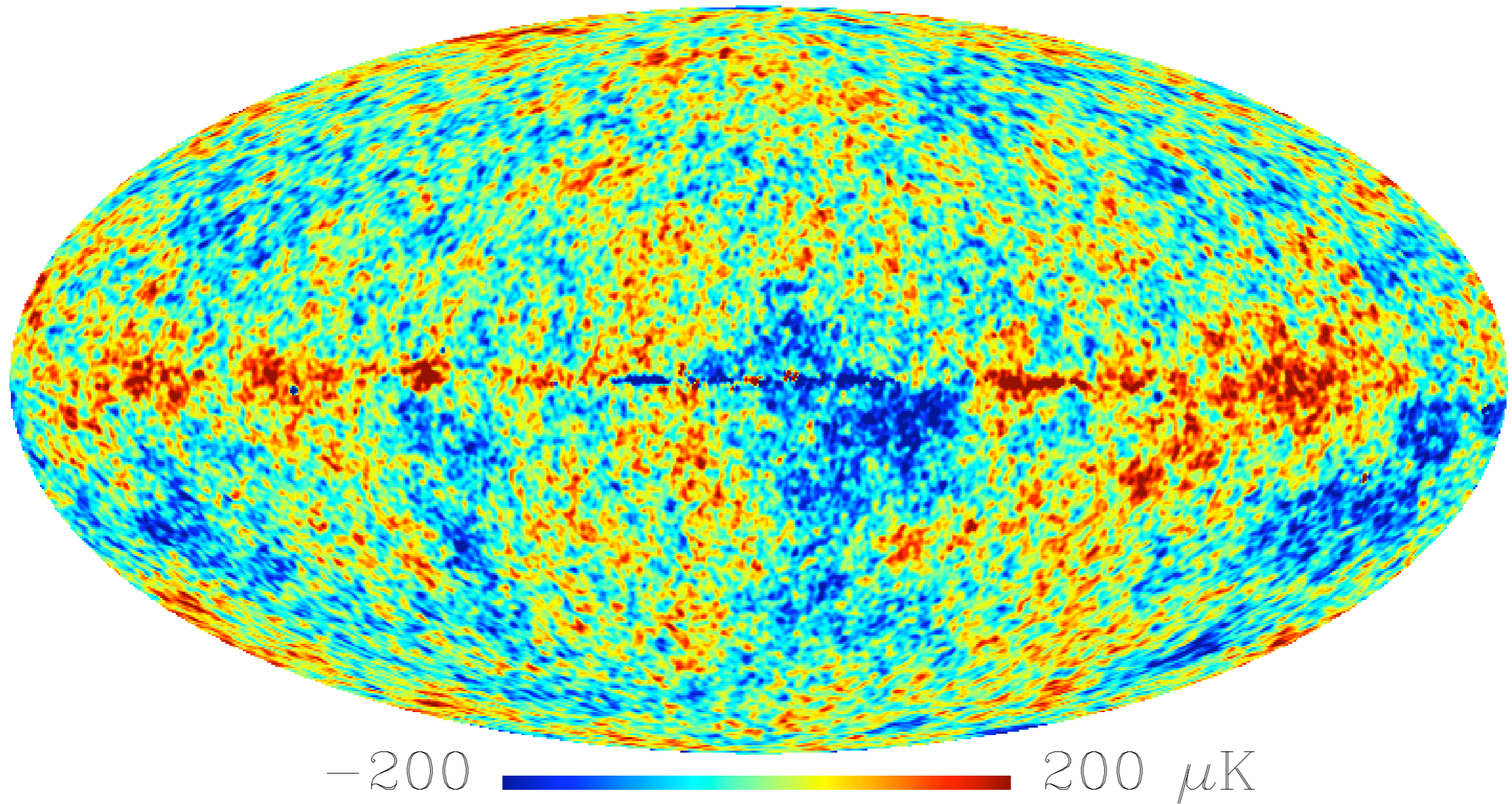
- public code: <http://research.ics.tkk.fi/ica/fastica/>
- demo: [http://research.ics.tkk.fi/ica/cocktail/cocktail\\_en.cgi](http://research.ics.tkk.fi/ica/cocktail/cocktail_en.cgi)
- different fields of application (not only astrophysics):
  - Impact of Contrast Functions in Fast-ICA on Twin ECG Separation (Fetal Electrocardiography) - Feb 2011
  - Energy-Efficient FastICA Implementation for Biomedical Signal Separation - Nov 2011
  - Separation of spinal cord motor signals using the FastICA method - Dec 2005
  - Speech Separation in the Vehicle Environment Based on FastICA Algorithm - Feb 2012
  - Palmprint recognition using FastICA algorithm and radial basis probabilistic neural network - Aug 2006
  - etc.....

# FastICA in astrophysics

- CMB
- 21-cm reionization signal
- Cosmological Stochastic Background of Gravitational Waves
- spectroscopic analysis of complex mixtures
- infrared emission spectra of interstellar dust
- .....



# FastICA on the CMB



application to WMAP 3yr data [Maino et al. MNRAS 374, 1207 (2007)]

# references

- FastICA technique:  
“Independent Component Analysis: A Tutorial”, A. Hyvarinen (1999) [Neural Networks, 13(4-5):411-430, 2000]
- FastICA in astrophysics:  
Maino et al., MNRAS, 334, 53 (2002)